

# Density-independent effective interaction for mean-field and beyond mean-field calculations

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From nuclear structure to particle transfer reactions and back,  
ETC\*, November 4-8, 2013



- “Standard” effective interactions and/or EDFs
- Extended and regularized Skyrme interactions
- Properties in infinite nuclear matter
- Preliminary results for finite nuclei

- Effective interactions and/or effective Energy Density Functionals

$$E = \langle \hat{T} + \hat{V}_{\text{eff}} \rangle = \int \mathcal{E}(\mathbf{r}) d^3r$$

are the key ingredients for the mean-field and beyond mean-field calculations

- They depend on coupling constants fitted on
  - infinite nuclear matter empirical properties
  - nuclear data (masses, radii, ...)
- Used at the mean-field level: “SR-EDF” (HF, HFB, ...)
- or beyond: RPA or “MR-EDF” (GCM, symmetry restauration)

# A brief history of the effective interactions

- Effective interactions in non-relativistic approaches

$$\hat{V}_{\text{eff}} \equiv V_{12} + V_{\text{so}} + V_{\text{Coul.}} + V_{\text{T}}$$

- Two families of (early) interactions

- Skyrme zero-range interaction

$$V_{12} \propto \delta(\mathbf{r}_2 - \mathbf{r}_1)$$

depends on gradients, gives a functional of local densities, needs a cut-off.

- Brink-Boeker (and Gogny) finite-range interaction

$$V_{12} \propto e^{-\frac{(\mathbf{r}_2 - \mathbf{r}_1)^2}{\mu^2}}$$

gives a functional of the non-local density, no cut-off.

## Need for a three-body interaction

- Very poor properties for the saturation point of infinite nuclear matter (“SV” interaction)

$$\rho_{\text{sat}} = 0.155 \text{ fm}^{-3}, \quad B = -16.05 \text{ MeV}, \quad K_{\infty} = 305.7 \text{ MeV},$$

$$m^*/m = 0.38, \quad J = 32.8 \text{ MeV}.$$

- Need for a three-body term

$$V_{123} = t_3 \delta(\mathbf{r}_2 - \mathbf{r}_1) \delta(\mathbf{r}_3 - \mathbf{r}_1).$$

→ improves the effective mass but...

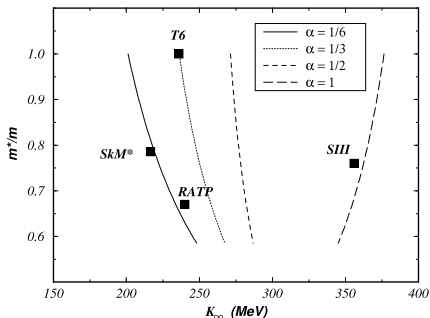
## From a three-body term to a density dependent term

- Collapse a spin polarized matter !!

$$t_3 \delta(\mathbf{r}_2 - \mathbf{r}_1) \delta(\mathbf{r}_3 - \mathbf{r}_1) \longrightarrow \frac{1}{6} t_3 \rho_0 \delta(\mathbf{r}_2 - \mathbf{r}_3).$$

not an interaction anymore.

- Incompressibility much to high !



$$\frac{1}{6} t_3 \rho_0 \delta(\mathbf{r}_2 - \mathbf{r}_3) \longrightarrow \frac{1}{6} t_3 \rho_0^\alpha \delta(\mathbf{r}_2 - \mathbf{r}_3).$$

# The standard (2-body) Skyrme functional

## ■ Effective Skyrme *interaction*

$$\begin{aligned} V_{\text{eff}} &= t_0 (1 + x_0 \hat{P}^\sigma) \delta && \text{local} \\ &+ \frac{t_1}{2} (1 + x_1 \hat{P}^\sigma) (\mathbf{k}'^2 \delta + \delta \mathbf{k}^2) && \text{non local} \\ &+ t_2 (1 + x_2 \hat{P}^\sigma) \mathbf{k}' \cdot \delta \mathbf{k} && \text{non local} \\ &+ \frac{t_3}{6} (1 + x_3 \hat{P}^\sigma) \rho_0^\alpha \delta && \text{density dep.} \\ &+ i W_0 \hat{\sigma} \cdot [\mathbf{k}' \times \delta \mathbf{k}] && \text{spin-orbit} \end{aligned}$$

- Sometimes complemented with tensor, D-wave terms, etc.
- Higher order derivatives ? Other density dependent terms ?
- Different interaction in the pairing channel
- $\rho_0^\alpha$  is the key to succes:
  - Incompressibility
  - Effective mass
  - Stability in the spin channels

# Effective interactions Vs. Functionals

## ■ Interactions:

- time-even, time-odd and pairing parts of the functional are entirely determined by the interaction parameters
- Difficult to have satisfying properties in all channels
- Few observables to constrain the time-odd terms

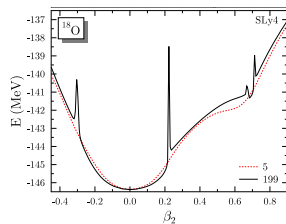
## ■ Functionals: more flexible

- Complicated, poorly determined or “dangerous” terms, *e.g.*  $\mathbb{J}^2$ ,  $\rho_1 \Delta \rho_1$ ,  $\mathbf{s}_0 \Delta \mathbf{s}_0$ ,  $\mathbf{s}_1 \Delta \mathbf{s}_1$ , ... can be separately adjusted or disregarded.
- Simpler interactions can be used in the pairing channel

⇒ Very efficient at the mean-field level



## ■ Beyond mean field calculations with a Skyrme EDF



(with terms  $\propto \rho^\alpha$ ,  $\alpha \notin \mathbb{N}$ )

Poles  $\rightarrow$  that can be corrected<sup>1</sup>  
and steps  $\rightarrow$  that can not !  
in the  
projected  
energy

... what about coul-ex ?

Introduction

Effective interactions / EDF

New Skyrme interaction

Regularized Skyrme interaction

Three-body term

Summary

See:

M. Anguiano *et al.* NPA 696 (2001) 467,  
J. Dobaczewski *et al.* PRC 76, 054315 (2007),  
PRC 79, 044318, 044319 and 044320,  
and L.M. Robledo, J. Phys. G 37, 064020 (2010).

- $\Rightarrow$  Need for an interaction (*i.e.* without  $\rho_0^\alpha$ ) for the Hartree, Fock and pairing terms !
- $\Rightarrow$  Three-body interaction ? (Cf J. Sadoudi thesis, CEA Saclay)
- $\Rightarrow$  Four-body interaction ? Something else ?

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<sup>1</sup>in some specific cases only !

## ■ Two-body effective interaction

$$\begin{aligned} V_{\text{eff}} &= t_0 (1 + x_0 \hat{P}^\sigma) \delta && \text{local} \\ &+ \frac{1}{2} t_1 (1 + x_1 \hat{P}^\sigma) (\mathbf{k}'^2 \delta + \delta \mathbf{k}^2) && \text{non local} \\ &+ t_2 (1 + x_2 \hat{P}^\sigma) \mathbf{k}' \cdot \delta \mathbf{k} && \text{non local} \\ &+ i W_0 \hat{\boldsymbol{\sigma}} \cdot [\mathbf{k}' \times \delta \mathbf{k}] && \text{spin-orbit} \end{aligned}$$

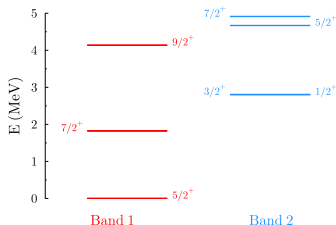
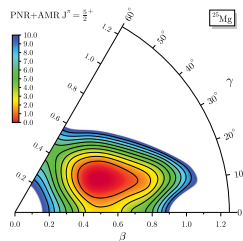
## ■ Complemented it with

$$\begin{aligned} &3 u_0 \delta_{12} \delta_{13} \\ &+ \frac{3}{2} u_1 (1 + y_1 \hat{P}^\sigma) \left[ \delta_{12} \delta_{13} \mathbf{k}_{12}^2 + \mathbf{k}_{12}'^2 \delta_{12} \delta_{13} \right] \\ &+ 3 u_2 (1 + y_{21} \hat{P}_{12}^\sigma + y_{22} \hat{P}_{13}^\sigma) \mathbf{k}_{12}' \cdot \delta_{12} \delta_{13} \mathbf{k}_{12} \\ &+ v_0 \delta_{12} \delta_{13} \delta_{14} \end{aligned}$$

## ■ And possibly: tensor, D-wave and 3-body spin-orbit...

First exploratory work: [J. Sadoudi \*et al.\*, Phys. Scr. 2013 014013.](#)

## Odd-even nucleus with symmetry restoration



A primer in nuclear structure !  
Calculation based on an interaction (SLyMR0)  
One quasi-particle state  
Projection on  $N$ ,  $Z$  and  $J$

Calculations by B. Bally and collaborators (CENBG, Bordeaux)

# Regularized Skyrme effective interaction

Density-independent interaction

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## ■ Interaction at NLO

$$\begin{aligned}v &= \tilde{v}_0(\mathbf{r}_1, \mathbf{r}_2; \mathbf{r}_3, \mathbf{r}_4) t_0 (1_{\sigma q} + x_0 1_q \hat{P}^\sigma - y_0 1_\sigma \hat{P}^q - z_0 \hat{P}^\sigma \hat{P}^q) \\ &+ \tilde{v}_1(\mathbf{r}_1, \mathbf{r}_2; \mathbf{r}_3, \mathbf{r}_4) t_1 (1_{\sigma q} + x_1 1_q \hat{P}^\sigma - y_1 1_\sigma \hat{P}^q - z_1 \hat{P}^\sigma \hat{P}^q) \\ &+ \tilde{v}_2(\mathbf{r}_1, \mathbf{r}_2; \mathbf{r}_3, \mathbf{r}_4) t_2 (1_{\sigma q} + x_2 1_q \hat{P}^\sigma - y_2 1_\sigma \hat{P}^q - z_2 \hat{P}^\sigma \hat{P}^q)\end{aligned}$$

Introduction

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New Skyrme interaction

Regularized Skyrme interaction

Three-body term

Summary

with

$$\tilde{v}_0(\mathbf{r}_1, \mathbf{r}_2; \mathbf{r}_3, \mathbf{r}_4) = \delta(\mathbf{r}_1 - \mathbf{r}_3) \delta(\mathbf{r}_2 - \mathbf{r}_4) g_a(\mathbf{r}_1 - \mathbf{r}_2)$$

$$\tilde{v}_1(\mathbf{r}_1, \mathbf{r}_2; \mathbf{r}_3, \mathbf{r}_4) = \delta(\mathbf{r}_1 - \mathbf{r}_3) \delta(\mathbf{r}_2 - \mathbf{r}_4) g_a(\mathbf{r}_1 - \mathbf{r}_2) \frac{1}{2} [\mathbf{k}_{12}^{*2} + \mathbf{k}_{34}^2]$$

$$\tilde{v}_2(\mathbf{r}_1, \mathbf{r}_2; \mathbf{r}_3, \mathbf{r}_4) = \delta(\mathbf{r}_1 - \mathbf{r}_3) \delta(\mathbf{r}_2 - \mathbf{r}_4) g_a(\mathbf{r}_1 - \mathbf{r}_2) \mathbf{k}_{12}^* \cdot \mathbf{k}_{34}$$

and

$$g_a(\mathbf{r}) = \frac{e^{-\frac{r^2}{a^2}}}{(a\sqrt{\pi})^3}$$

- Higher order derivative terms provide an order-by-order correctible theory

See J. Dobaczewski, K.B., F. Raimondi, J. Phys. G **39**, 125103

## Preliminary remarks

- For the moment: Standard zero-range spin-orbit interaction
- Two flavors for this interaction
  - $t_2 = -t_1$ : “REG2a” interaction  $\rightarrow$  much simpler...
  - $t_2 \neq -t_1$ : “REG2b” interaction
- **Two-body** effective interaction
- **No** density dependence !
- Simple expression with 12 parameters (plus the spin-orbit term)
- Compact expressions for infinite nuclear matter properties (Cf. arXiv:1305.7210)
- Relatively simple to implement in a code on harmonic oscillator basis (done in HFODD)
  
- Is it possible to reproduce the empirical properties of the saturation point ?
- Is it possible to have correct binding energies for nuclei ?

For  $i = 0, 1$  and 2: 4 direct + 4 exch. terms with coupling constants

$$\begin{cases} A_i^{\rho 0} &= \frac{1}{2} t_i \left( 1 + \frac{1}{2} x_i - \frac{1}{2} y_i - \frac{1}{4} z_i \right) \\ A_i^{\rho 1} &= -\frac{1}{2} t_i \left( \frac{1}{2} y_i + \frac{1}{4} z_i \right) \\ A_i^{\text{s}0} &= \frac{1}{2} t_i \left( \frac{1}{2} x_i - \frac{1}{4} z_i \right) \\ A_i^{\text{s}1} &= -\frac{1}{8} t_i z_i \end{cases}$$

and

$$\begin{cases} B_i^{\rho 0} &= -\frac{1}{2} t_i \left( \frac{1}{4} + \frac{1}{2} x_i - \frac{1}{2} y_i - z_i \right) \\ B_i^{\rho 1} &= -\frac{1}{2} t_i \left( \frac{1}{4} + \frac{1}{2} x_i \right) \\ B_i^{\text{s}0} &= -\frac{1}{2} t_i \left( \frac{1}{4} - \frac{1}{2} y_i \right) \\ B_i^{\text{s}1} &= -\frac{1}{8} t_i \end{cases}$$

- 8 independent coupling constants for REG2a, 12 for REG2b
- More degrees of freedom to adjust the different channels than with a zero-range interaction

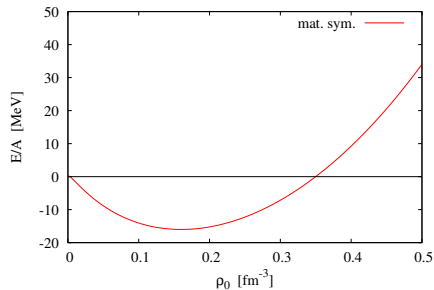
# Adjustment of the parameters

- No spherical code (for now)...
- HFODD: 1 magic nucleus on 10 shells  $\simeq$  15 minutes
- Use of infinite nuclear matter to limit the number of free parameters
- Handmade non professional adjustment (*POUNDerS* on the way)
- Exact treatment of the Coulomb interaction

# Infinite nuclear matter (REG2b)

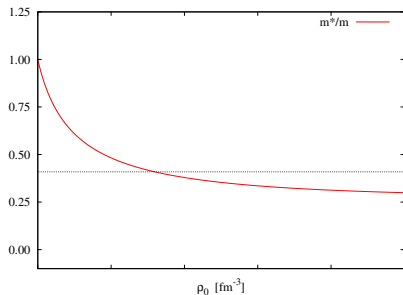
Density-independent interaction

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$$a = 0.8 \text{ fm}$$

$$\begin{aligned} \rho_{\text{sat}} &= 0.16 \text{ fm}^{-3} \\ E/A &= -16 \text{ MeV} \\ K_{\infty} &= 230 \text{ MeV} \\ J & \\ L & \\ K_{\text{sym}} & \end{aligned}$$



$$m^*/m = 0.41 \quad :-($$

masses of doubly magic nuclei

Introduction

Effective interactions / EDF

New Skyrme interaction

Regularized Skyrme interaction

Three-body term

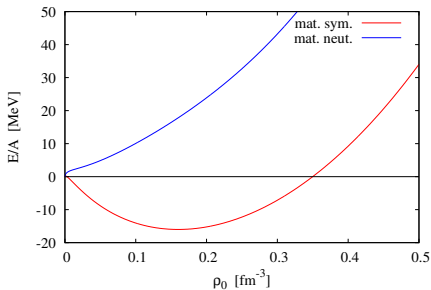
Summary



# Infinite nuclear matter (REG2b)

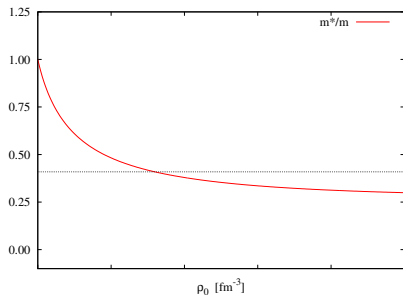
Density-independent interaction

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$$a = 0.8 \text{ fm}$$

$$\begin{aligned} \rho_{\text{sat}} &= 0.16 \text{ fm}^{-3} \\ E/A &= -16 \text{ MeV} \\ K_{\infty} &= 230 \text{ MeV} \\ J &= 32 \text{ MeV} \\ L &= 58 \text{ MeV} \\ K_{\text{sym}} &= -175 \text{ MeV} \end{aligned}$$



$$m^*/m = 0.41 \text{ :-(}$$

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Introduction

Effective interactions / EDF

New Skyrme interaction

Regularized Skyrme interaction

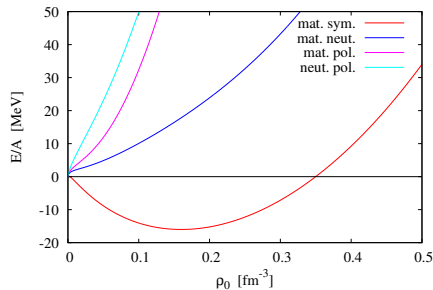
Three-body term

Summary

# Infinite nuclear matter (REG2b)

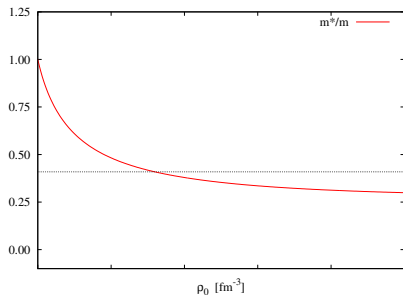
Density-independent interaction

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$$a = 0.8 \text{ fm}$$

$\rho_{\text{sat}}$	=	$0.16 \text{ fm}^{-3}$
$E/A$	=	$-16 \text{ MeV}$
$K_{\infty}$	=	$230 \text{ MeV}$
$J$	=	$32 \text{ MeV}$
$L$	=	$58 \text{ MeV}$
$K_{\text{sym}}$	=	$-175 \text{ MeV}$



$$m^*/m = 0.41 \quad :-($$

masses of doubly  
magic nuclei

Introduction

Effective interactions / EDF

New Skyrme interaction

Regularized Skyrme interaction

Three-body term

Summary

# Comparison in infinite nuclear matter

Density-independent interaction

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Introduction

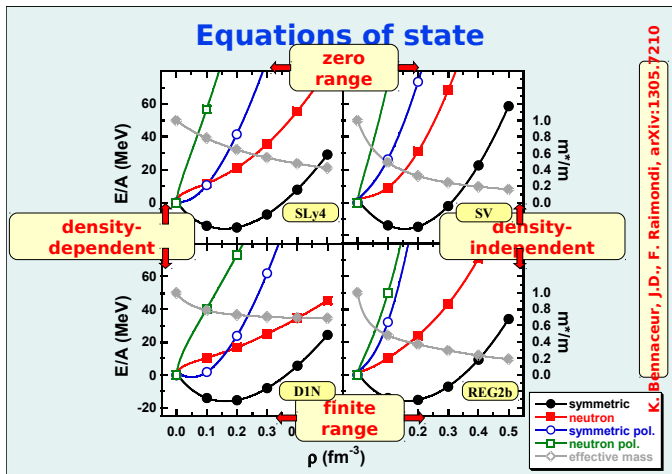
Effective interactions / EDF

New Skyrme interaction

Regularized Skyrme interaction

Three-body term

Summary



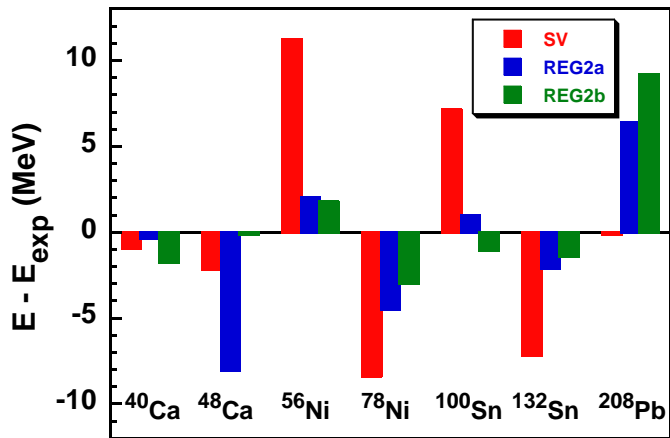
Jacek Dobaczewski



JYVÄSKYLÄN YLIOPISTO



## Masses of doubly magic nuclei



Better than SV, not a tremendous achievement.

How to correct the effective mass ?

- One can consider a finite range 3-body interaction

$$V_{3N} = C_{3N} \times e^{-\frac{(r_2-r_1)^2+(r_3-r_2)^2+(r_1-r_3)^2}{\alpha_{3N}^2}}$$

as suggested by A. Zapp, R. Roth, H. Hergert<sup>2</sup>

- Or an easier the handle 3-body zero-range

$$v_3 = t_3 \delta(\mathbf{r}_2 - \mathbf{r}_1) \delta(\mathbf{r}_3 - \mathbf{r}_1)$$

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<sup>2</sup>unpublished, as far as I know...

## Zero-range 2-body

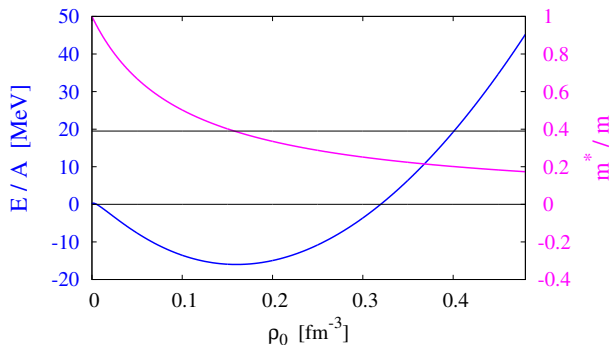
- Two-body zero-range  $\simeq$  SV interaction

- Fixed:

$$\rho_{\text{sat}} = 0.16 \text{ fm}^{-3}, \quad B = -16 \text{ MeV}$$

- Calculated:

$$K_{\infty} = 306 \text{ MeV}, \quad Q = -174 \text{ MeV}, \quad m^*/m = 0.39$$



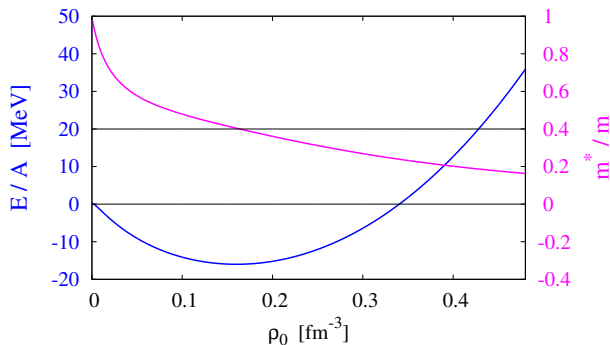
## Finite-range 2-body

- Finite-range 2-body ( $a = 1$  fm)  $\simeq$  Reg2b interaction
- Fixed:

$$\begin{aligned}\rho_{\text{sat}} &= 0.16 \text{ fm}^{-3}, & K_{\infty} &= 230 \text{ MeV}, \\ B &= -16 \text{ MeV}, & Q &= -174 \text{ MeV}\end{aligned}$$

- Calculated:

$$m^*/m = 0.40$$



# Finite-range 2-body + Zero-range 3-body

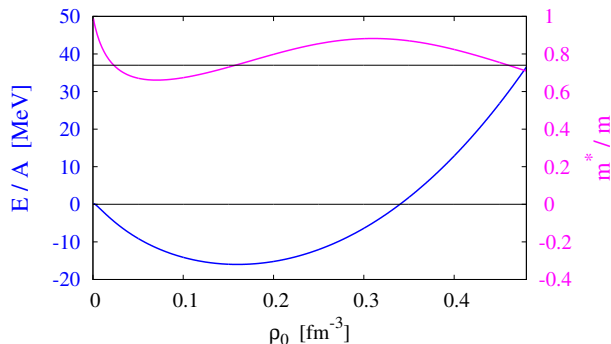
- Finite-range 2-body ( $a = 1$  fm),  
 $t_3 = 9600$  MeV fm<sup>6</sup> adjusted by hand

- Fixed:

$$\begin{aligned} \rho_{\text{sat}} &= 0.16 \text{ fm}^{-3}, & K_{\infty} &= 230 \text{ MeV}, \\ B &= -16 \text{ MeV}, & Q &= -174 \text{ MeV} \end{aligned}$$

- Calculated:

$$m^*/m = 0.74$$





# Finite-range 2-body + Zero-range 3-body

Density-independent interaction

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- First exploratory work:
  - “Reg2a” form + Zero-range 3-body
  - parameters adjusted so that masses of doubly magic nuclei are correct within 5 %
- Interaction stable in all channels (comparison with D1N)

Introduction

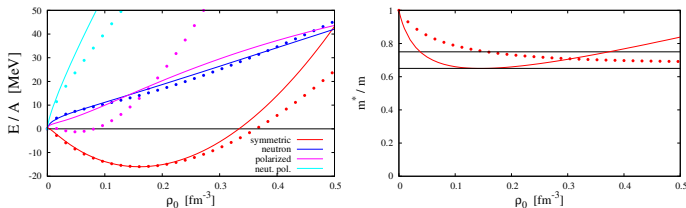
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$$\rho_{\text{sat}} = 0.16 \text{ fm}^{-3}, \quad B = -16 \text{ MeV}, \quad K_{\infty} = 265 \text{ MeV}$$
$$m^*/m = 0.65, \quad J = 31.3 \text{ MeV}$$

- Ready for a fit...

- The final goal of this work is to develop an effective interaction that can be used in **mean-field** and **beyond mean-field** calculations
- The 2-body version reproduce correctly the empirical saturation properties of infinite nuclear matter but the **effective mass** that is **too low**
- A **3-body contact** term seems to solve this problem
- Properties in the **pairing channel** not studied yet...

## Work done in collaboration with

- J. Dobaczewski (Univ. of Warsaw / Univ. of Jyväskylä)
  - F. Raimondi (TRIUMF)
- 



Thank  
you...



Density-  
independent  
interaction

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Introduction

Effective  
interactions /  
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New Skyrme  
interaction

Regularized  
Skyrme  
interaction

Three-body term

Summary